Turning on the cosmic microphone

A new tool lets astronomers 'listen' to the Universe for the first time.

By Diana Kwon

On the morning of 14 September 2015, a signal from two black holes that had collided 1.3 billion years ago reached Earth, alerting scientists around the world. "It took us a good part of the day to convince ourselves that this was not a drill", says Professor Gabriela González. In fact, it was the first-ever detection of a gravitational wave and the latest development in the long history of astronomy.

When Galileo first introduced the telescope in the 1600s, astronomers gained the ability to view parts of the Universe that were invisible to the naked eye. This led to centuries of discovery – as telescopes advanced, they exposed new planets, galaxies and even a glimpse of the very early Universe. In 2015, scientists gained another valulable tool: the ability to 'hear' the cosmos through gravitational waves.

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REVIEW

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Great news for us all – after years of looking for gravitational waves, scientists have finally detected them. This article makes good reading for physics and general science teachers. It can be used as a platform for discussion in class, focusing mainly on the problems associated with detecting gravitational waves, but more importantly on their applications in everyday life.

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A simulation of two black holes merging and the resulting emission of gravitational radiation

Image courtesy of LIGO / Penn State; image source: Flickr



Aerial view of the LIGO gravitational-wave detector in Livingston, Louisiana

Ripples in space-time

Newton described gravity as a force. Thinking about gravity in this way can explain most of the phenomena that happen here on Earth. For example, the force of gravity acting on an apple makes it fall from a tree onto an unsuspecting person sitting below it. However, to understand gravity on a cosmic scale, we need to turn to Einstein, who described gravity as the bending of space-time itself.

Some physicists describe this process using a bowling ball and a blanket. Imagine space-time as a blanket. A bowling ball placed at the centre of the blanket bends the fabric around it. The heavier an object is, the further it sinks. As you move the ball along the fabric, it produces ripples, much like a boat travelling through water.

"The curvature is what makes Earth orbit the Sun – the Sun is a bowling ball in a fabric and it's that bending in the fabric that makes the Earth go around", explains González, who is the spokesperson for the Laser Interferometer Gravitational-Wave Observatory (LIGO) collaboration.

Everything that has mass – planets, stars and people – pulls on the fabric of space-time and produces gravitational waves as it moves through space. These waves are passing through us all the time, but they are much too weak to detect.

To find these elusive signals, physicists built LIGO, twin observatories in Louisiana and Washington, USA. At each L-shaped detector, a laser beam is split and sent down two 4 km arms. The beams reflect off the mirrors at each end and travel back to reunite. A passing gravitational wave slightly alters the relative lengths of the arms, shifting the path of the laser beam, creating a change that physicists can detect.

Unlike telescopes, which are pointed toward very specific parts of the sky, detectors like LIGO scan a much larger area of the Universe and hear sources from all directions. "Gravitational waves



detectors are like microphones", says Laura Nuttall, a post-doctoral researcher at Syracuse University, USA.

First detections

On that September morning in 2015, when the first gravitational wave passed through the two detectors, LIGO was still preparing for an observational run. Researchers were still running tests and diagnostics during the day – which is why they needed to conduct a large number of checks and analyses to make sure the signal was real.

Months later, once researchers had meticulously checked the data for errors or noise (such as lightning or earthquakes), the LIGO collaboration announced to the world that they had finally reached a long-anticipated goal: almost 100 years after Einstein first predicted the existence of gravitation waves, scientists had detected them.

A few months after the first signal arrived, LIGO detected yet another

black-hole collision. "Finding a second one proves that there's a population of sources that will produce detectable gravitational waves", Nuttall says. "We are actually an observatory now."

Cosmic microphones

Many have dubbed the detection as the dawn of the age of gravitational-wave astronomy. Scientists expect to see hundreds, maybe even thousands, of these binary black holes in the years to come. Gravitational-wave detectors will also allow astronomers to look much more closely at other astronomical phenomena, such as neutron stars, supernovae and even the Big Bang.

One important next step is to detect the optical counterparts – such as light from the surrounding matter or gamma-ray bursts – of the sources of gravitational waves. To do this, astronomers need to point their telescopes to the area of the sky where the gravitational waves came from to find any detectable light.

Currently, this feat is like finding a needle in a haystack. Because the field of view of gravitational-wave detectors is much, much larger than telescopes, it is extremely difficult to connect the two. "Connecting gravitational waves with light for the first time will be such an important discovery that it's definitely worth the effort", says Edo Berger, an astronomy professor at Harvard University.

LIGO is only one of several gravitational-wave observatories. Other ground-based observatories, such as Virgo in Italy, KAGRA in Japan and the future LIGO India have similar sensitivities to LIGO. There are also other approaches that scientists are using – and plan to use in the future – to detect gravitational waves at completely different frequencies.

The evolved Laser Interferometer Space Antenna (eLISA), for example, is a gravitational-wave detector that physicists plan to build in space. Once complete, eLISA will be composed of

Image courtesy of Penn State; image source: Flickr



three spacecraft that are over a million kilometres apart, making it sensitive to much lower gravitational-wave frequencies, where scientists expect to detect supermassive black holes.

Pulsar array timing is a completely different method of detection. Pulsars are natural timekeepers, regularly emitting beams of electromagnetic radiation. Astronomers carefully measure the arrival time of the pulses to find discrepancies, because when a gravitational wave passes by, spacetime warps, changing the distance between us and the pulsar, causing the pulses to arrive slightly earlier or later. This method is sensitive to even lower frequencies than eLISA can detect.

These and many other observatories will reveal a new view of the Universe, helping scientists to study phenomena such as merging black holes, to test theories of gravity and possibly even to discover something completely unexpected, says Daniel Holz, a professor of physics and astronomy at the University of Chicago. "Usually in science you're just pushing the boundaries a little bit, but in this case, we're opening up a whole new frontier."

Diana is a freelance science journalist based in Berlin, Germany. Her work has appeared both in print and online in numerous outlets including *Scientific American*, *Quartz* and *New Scientist*.

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